

# Nano-to-micro dust environment monitored by GIADA during the entire ROSETTA scientific phase

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The Micro Balance System (MBS) [1] is the GIADA [2,3,4] subsystem devoted to the flux and fluence of nano-to-micro dust particles measurement. The cumulative flux of particles/grains with diameters  $<10 \mu\text{m}$  is measured by a net of five Quartz Crystal Microbalances (QCMs) pointing towards different directions in order to characterize the dust flux within a solid angle of 180 deg.

Each QCM has an acceptance angle of about 40 deg, a collection area of about 12 mm<sup>2</sup> and consists of a matched pair of quartz crystals resonating at  $\sim 15\text{MHz}$ . Each QCM is equipped with a heating device to: (1) check the frequency vs. temperature dependence, (2) perform thermo-gravimetric measurements on the accumulated dust, at temperatures  $< 100 \text{ }^\circ\text{C}$ , and (3) remove volatile materials from the sensitive surface. Starting from the July 2014, i.e. during the Rosetta/ESA space probe approach to comet 67P/Churyumov-Gerasimenko, the MBS was continuously operating to monitor the dust coma environment. The QCMs' high sensitivity ( $0.2 \text{ [Hz ng}^{-1}\text{]}$ ) allowed to detect nano-micron-sized dust flux variation events well constrained in time. Otherwise, the nano-to-micron-sized particle flux was constant over the entire Rosetta mission scientific phase. The MBS data analysis allowed us to characterize the nano-to-micron sized dust particles flux identifying: (1) the preferred dust flux directions; (2) the flux time variation for particles with sizes smaller than 10 microns; 3) the presence of fine dust in dust "outbursts".

## References

[1] E.Palomba, E.L.Colangeli, P.Palumbo, A.Rotundi, J.M.Perrin, E.Bussoletti, Performance of micro-balances for dust flux measurement, *Adv.Space Res.* 29 (2002) 1155–1158, [http://dx.doi.org/10.1016/S0273-1177\(02\)00131-X](http://dx.doi.org/10.1016/S0273-1177(02)00131-X).

[2] V.DellaCorte, A.Rotundi, et al, GIADA: its status after the Rosetta cruise phase and on-ground activity in support of the encounter with comet 67P/Churyumov-Gerasimenko, *J.Astron.Instrum.* 3 (2014) 50011, <http://dx.doi.org/10.1142/S2251171713500116>.

[3] L.Colangeli, J.J.LopezMoreno, P.Palumbo, J.Rodriguez, E.Bussoletti, V.Della Corte, F.Esposito, M.Herranz, J.M.Jerónimo, A.Lopez-Jimenez, E.M.Epifani, R. Morales, E.Palomba, A.Rotundi & GIADA Team, GIADA: The grain impact analyser and dust accumulator for the Rosetta space mission, *Adv.Space Res.* 39(2007)446–450, <http://dx.doi.org/10.1016/j.asr.2006.12.048>.