



From LWS to ILWS and Norwegian Priorities

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ILWS Requirements

- **ILWS may have broader goals, motivation and requirements than LWS.**
- **Predictive capability still too small so scientific understanding must be main driver.**
- **Application side has proven useful to get funding for LWS, should not be oversold for ILWS.**
- **Coordination and possibly integration required with existing international “networking”.**



What is ILWS?

- **Core space infrastructure from major agencies.**
- **Add on from other agencies and institutions**
 - Payload provision
 - Technology development
 - Theory and modelling
 - Supplementary missions
 - Other observational tools (ground based, rockets, archives)
 - Education and outreach
 - Service providers



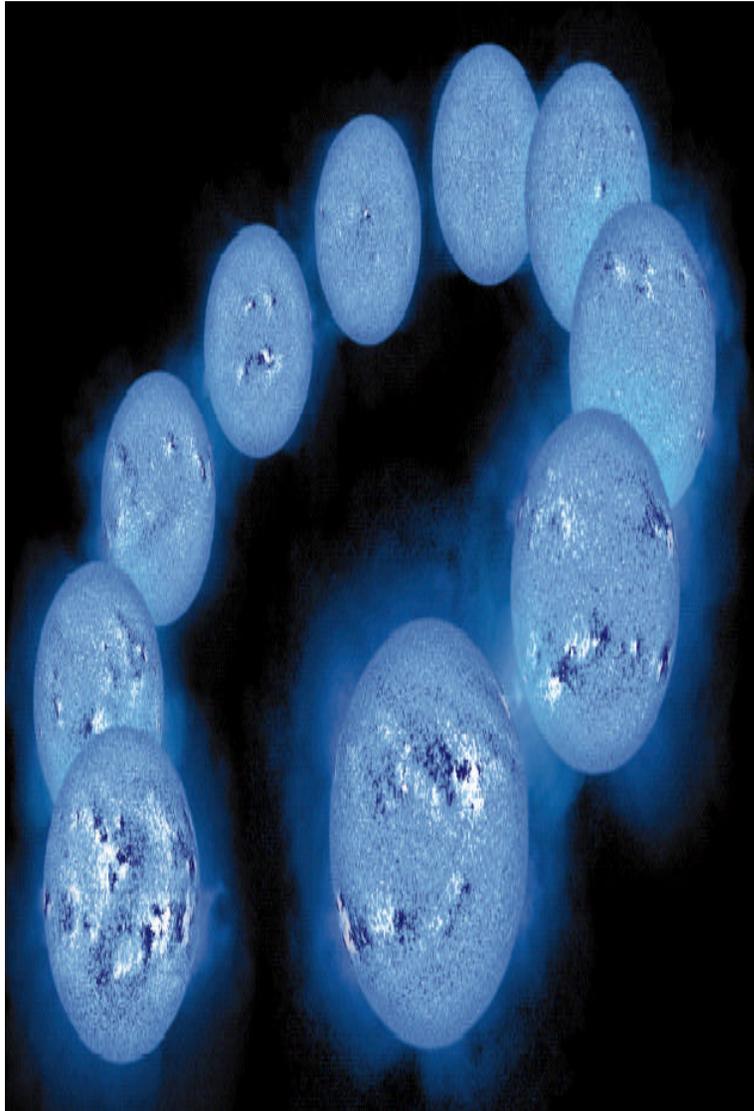
Norwegian Priorities (1)



- **Historical tradition in studies of Sun-Earth relationships (Birkeland & Størmer).**
- **Build up of solar physics from the 1950-ties.**
- **Scientific infrastructure, ARR, ALOMAR, SvalRak, EISCAT.**
- **Participation in several recent space missions: POLAR, Geotail, SOHO, Cluster.**
- **>85% of Norwegian space science is within solar-terrestrial physics.**



Norwegian Priorities (2)

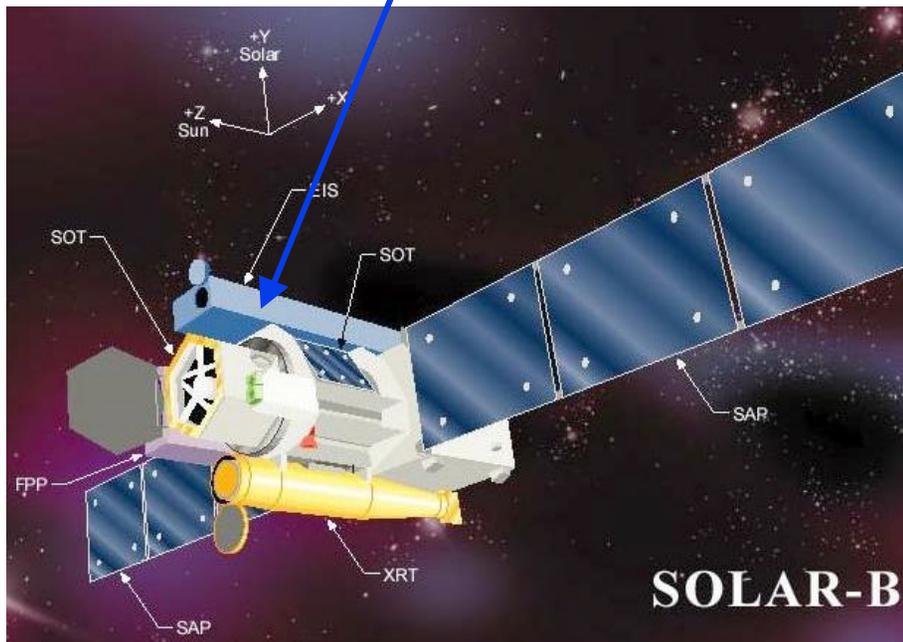


- To gain insight and understanding of the underlying physical processes governing:
 - The heating of the upper solar atmosphere.
 - The excited and structure of solar p-modes.
 - The microphysics of the auroral zone and polar cusp
 - Middle atmosphere dynamics
- Use of geographical advantage, the ESA membership and other suitable bilateral partnerships.
- National financing make pure non-ESA satellite projects difficult.



Solar Physics-Securing Continuity

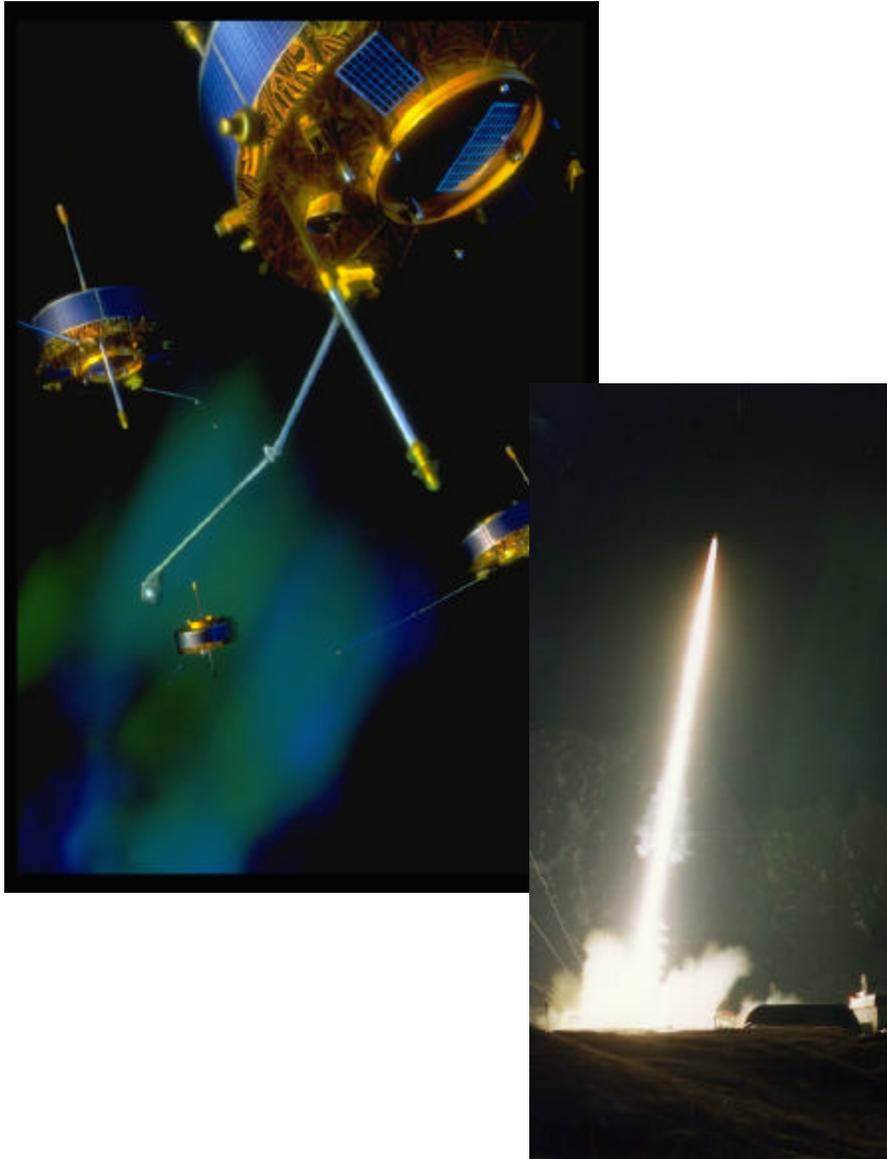
UiO with MSSL



- Clear line in priorities:
 - HRTS (rockets/shuttle)
 - SOHO
 - Solar B
 - SDO
- Strong in solar physics due to clear will to prioritize.
- Major setback for the future of solar physics in Norway with the non-selection of the two SDO proposals.
- Solar B and Solar Orbiter can only partially remedy this problem.
- More theory/modelling is needed to ensure the predictive power of an ILWS programme.



Space Physics-Complementary Tools



- **Combination of satellites, sounding rockets and ground measurements to gain added value.**
- **Strong participation i Cluster.**
- **Continuity required!**
- **For internationalization to ILWS it is necessary to look at both the STP and LWS missions.**
- **MMS, GEC as well as MC missions of interest to Norway.**



Leading Infrastructure



- **Andøya Rocket Range (69°N)**, more than 700 launches since 1962. 14 in ten hours this summer. The main polar launch site.
- **SvalRak (79°N)**, unique location in the polar cusp.
- **EISCAT** with sites on Svalbard in Norway, Sweden and Finland, both in the auroral zone and in the polar cusp.
- **SvalSat** is becoming the leading ground station for polar orbiting satellites.





New Approach to Rockets

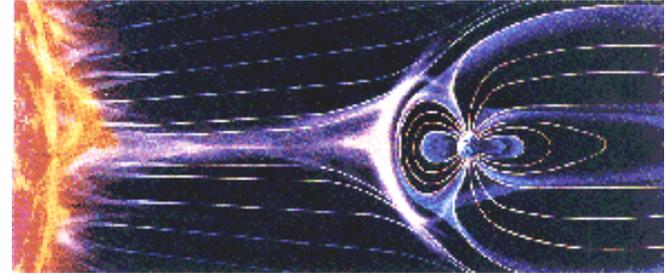


- Do only what satellites or ground based instrumentation cannot do:
 - High spatial and time resolution, and
 - Close links with other instrumentation, and
 - Low DV, specific conditions and directions, or
 - Altitude below 130 km
- Lower costs to get more launches and data:
 - Extremely small payloads (8 kg, 50 mm Ø, 15-20 k€) up to 110 km.
 - Recovery of medium sized and large expensive payloads up to 150 km
 - Use of standardized payload module and support infrastructure (Hotel Payload)
 - Increasing use for education and outreach



Svalbard- A Unique Location

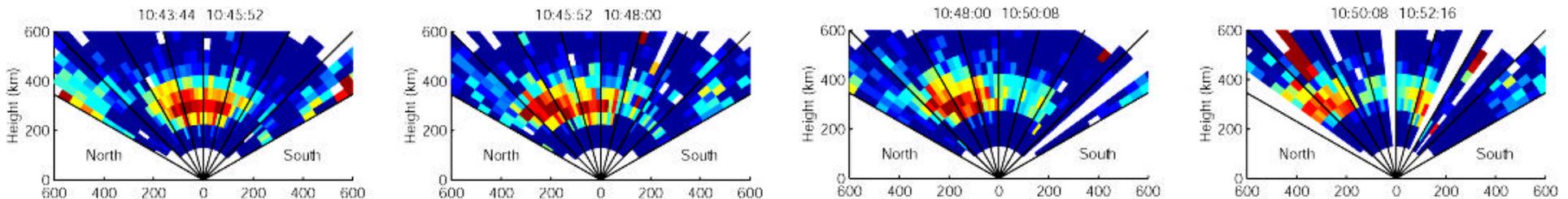
- Ideal location for daytime aurora
- Easy access, moderate climate and good general infrastructure
- Extended suite of scientific infrastructure
- Data reception from all orbits of polar orbiting satellites.





Plasma clouds with EISCAT

- Plasma clouds are created near the daytime aurora and drifts over the pole.
- EISCAT Svalbard Radar is a versatile tool to study the "stormy" surroundings of the aurora - the wind and density of the electron plasma.

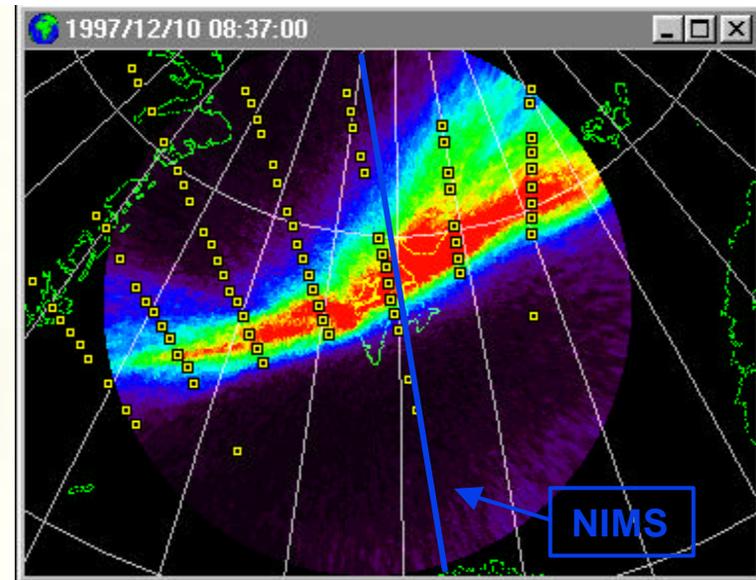
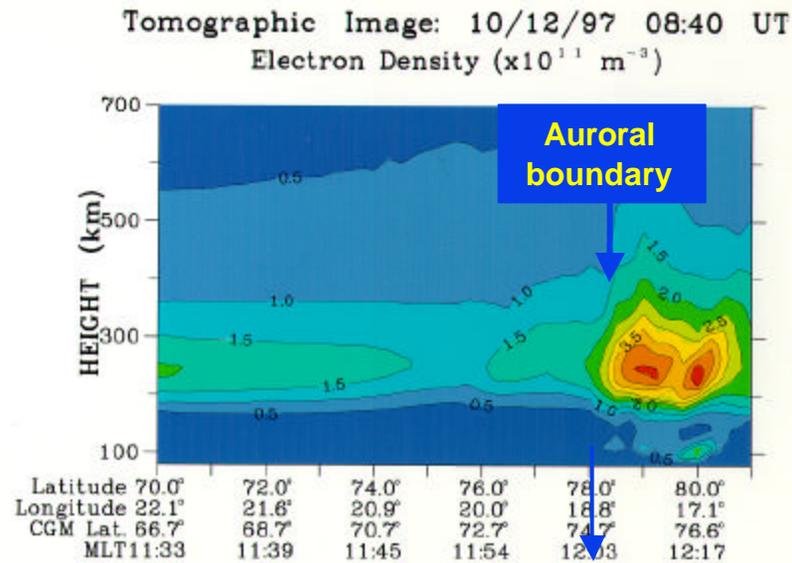


A big plasma cloud drifting northward on 19 December, 2001. In the rightmost picture we see a new cloud forming south of Svalbard.



Unresolved problem

What causes decametre scale plasma irregularities in the cusp for the HF radar to backscatter from?

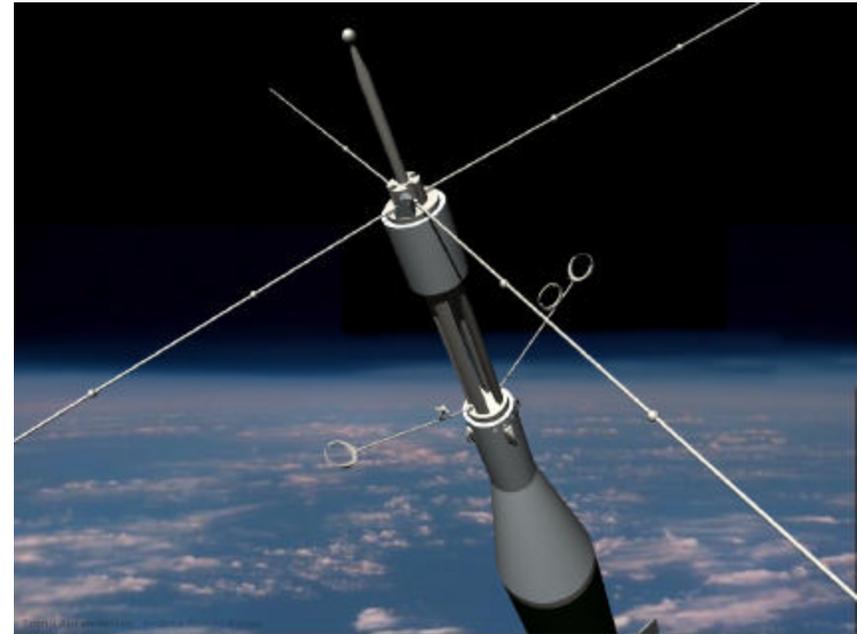
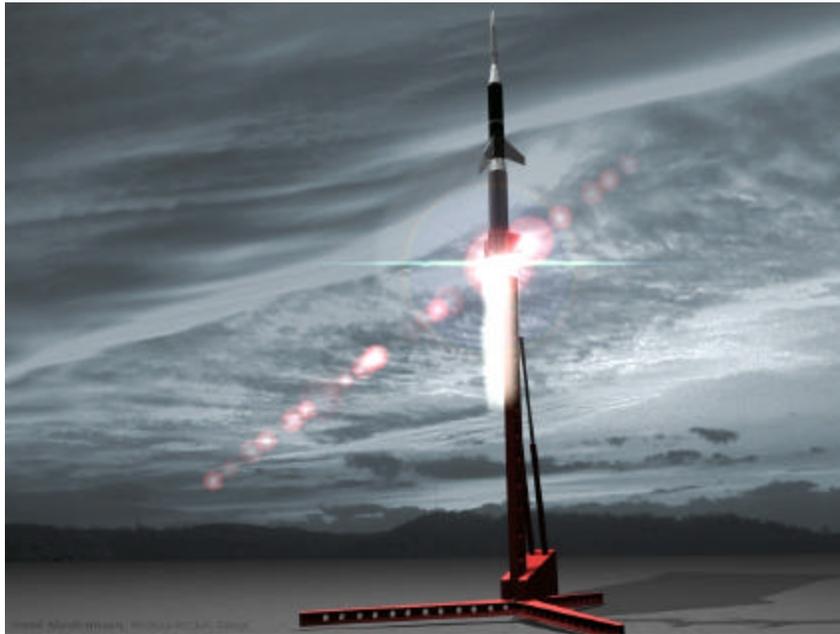


Gradient drift instability
growth rate

$$g = - \left(\frac{1}{n_0} \right) \left(\frac{\Delta n}{\Delta y} \right) \left(\frac{E}{B} \right) \implies 5-10 \text{ min}$$



ICI-1 Investigation of Cusp Irregularities





ICI-1

- **Electrical fields (DC/AC): UiO**
- **Positive ionprobe: FFI**
- **High energy particles: UiB**
- **Magnetic fields : CETP, Velizy**
- **Current measurements: LPCE, Orléans**
- **Payload: 50 kg**
- **Configuration: Nike/ Improved Orion**
- **Apogee: ~320 km**
- **December 2003**



Conclusions

- **Norway is strongly supportive of an ILWS programme with a primary science and a secondary application justification.**
- **Need a broad band of observational and theoretical tools.**
- **Unique contributions from Norway in fields of sounding rockets and infrastructure.**
- **On satellite side ESA connection is required for Norwegian national funding.**
- **Requires close co-ordination in mission planning and payload selection, specially between the major agencies.**