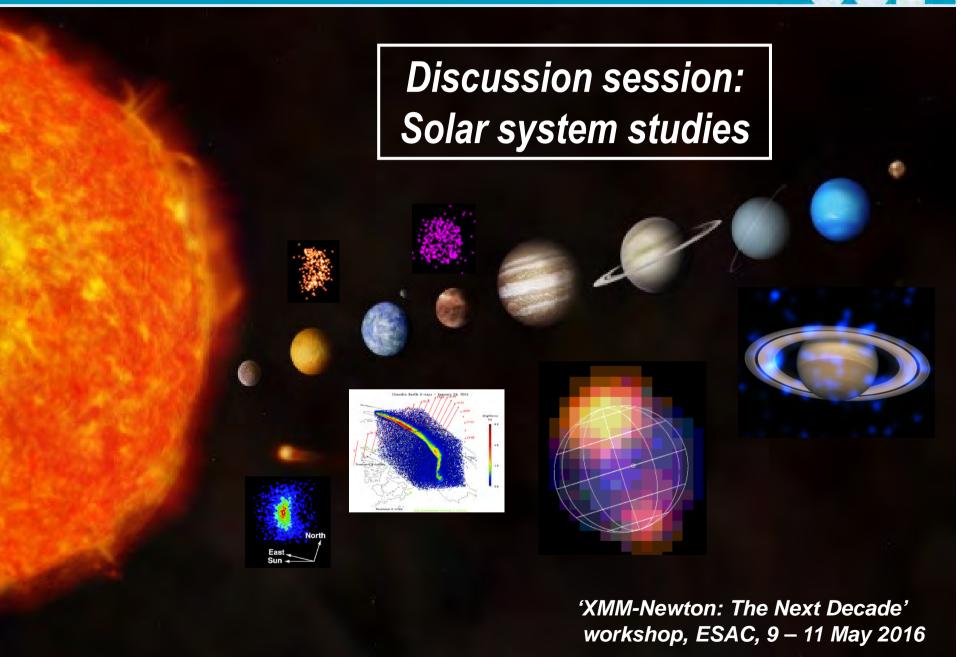
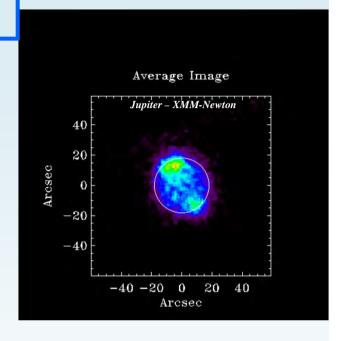
## **UCL**





### Looking ahead with XMM-Newton ...

- XMM-Newton and Chandra have demonstrated the potential of planetary X-ray astronomy
- XMM-Newton observations during enhanced solar activity & simultaneous with other facilities (Chandra, HST, Gemini, etc.) provide additional unparalleled scientific return
- Must exploit synergy with in-situ missions (like JUNO, 2016-17, and SMILE, 2021 onwards)

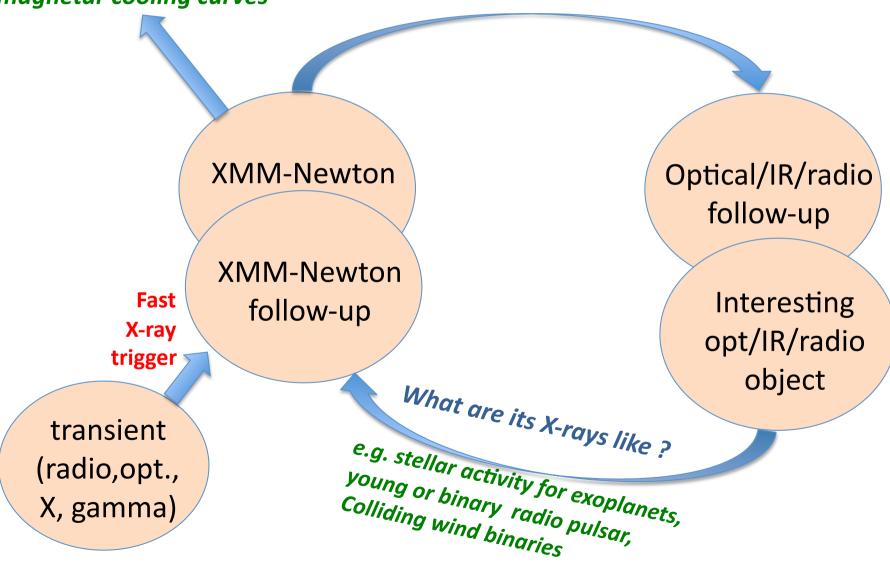


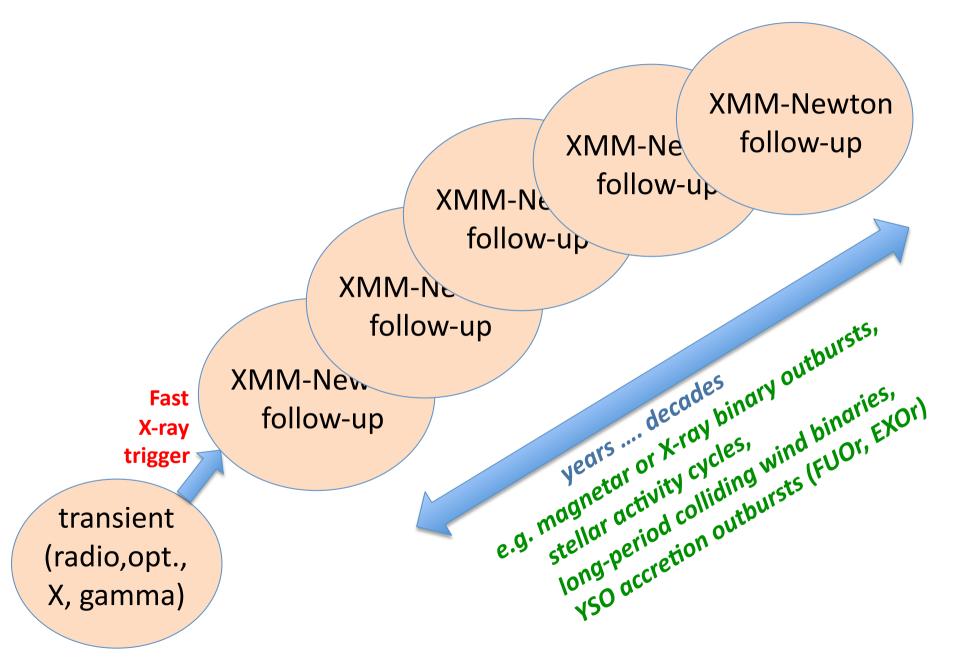
- Major objective is to determine how solar wind interacts with planetary magnetospheres (e.g. Jupiter) and exospheres (Mars, & Earth too!) and comets
- XMM-Newton by far the best mission for solar system X-ray studies:
   EPIC large FOV + unique combination of high sensitivity (especially at low energy) with moderate (EPIC) and superb (RGS) energy resolution, and OM (UV) particularly important for comets

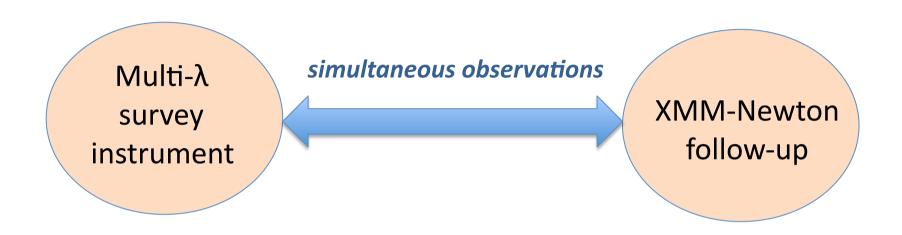
#### Great new physics

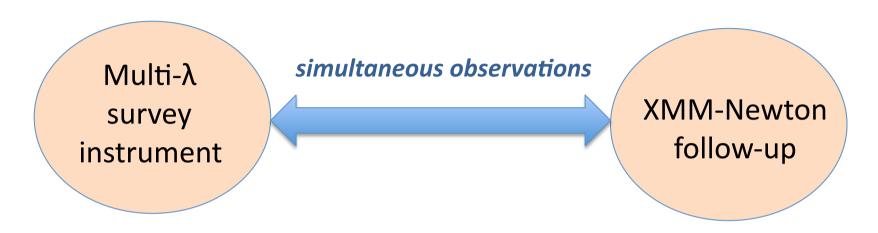
e.g. on outer stellar atmosphere, magnetar cooling curves

What is this X-ray source?
e.g. star vs AGN, stellar SpT, magnetar, XRB









#### e.g.

- stellar flares + rotation with photometric monitoring from space (K2, TESS, PLATO)
- stellar corona vs aurora emission with JVLA, SKA-precursors
- new millisecond pulsars from radio surveys (Parkes, GBT, LOFAR, ASKAP, MeerKAT)
- new transients in the X-ray all-sky monitors (Swift, Fermi-GBM, MAXI, eROSITA)
- new  $\gamma$  Cas analogs or colliding wind binary candidates (eROSITA)

#### XMM-Newton in the Next Decade

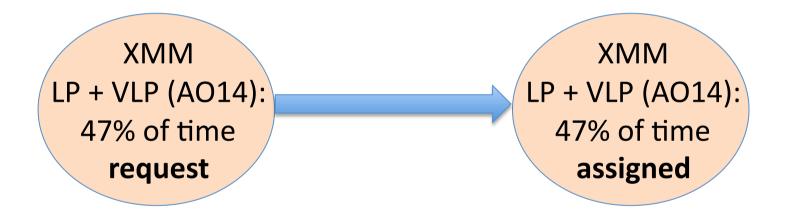
Which use of XMM-Newton will be most important in the Next Decade?

- (A) "stand-alone" science ( no need of other facilities )
- (B) simultaneous observations (opt/IR/radio + XMM-Newton)
- (C) "follow-up" science (opt/IR/radio  $\rightarrow$  XMM-Newton and vice versa)
- (D) fast triggered observations of Galactic and extra-galactic transients

# Need for longer and longer (nearly) uninterrupted observations

#### e.g.

- YSO accretion events triggered by flares (1-2 days)
- rotational modulation in winds of massive stars (few days to a week)
- rotational/orbital accretion variability (of flux and emission lines): few days



Is a change in policy needed to facilitate new science only accessible with long data sets?

## XMM-Newton strategy for "small" samples

