

Long-term transient alerts

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INTRODUCTION

Long-term variability

- Already >17200 observations with XMM-Newton
- Sources vary on the long term, no systematic study made in X-ray
- Some sources already detected >90 times
- Some sources observed and not detected
- Time domain astronomy becoming important
- XMM-Newton could continue until 2030s and could provide transient alerts

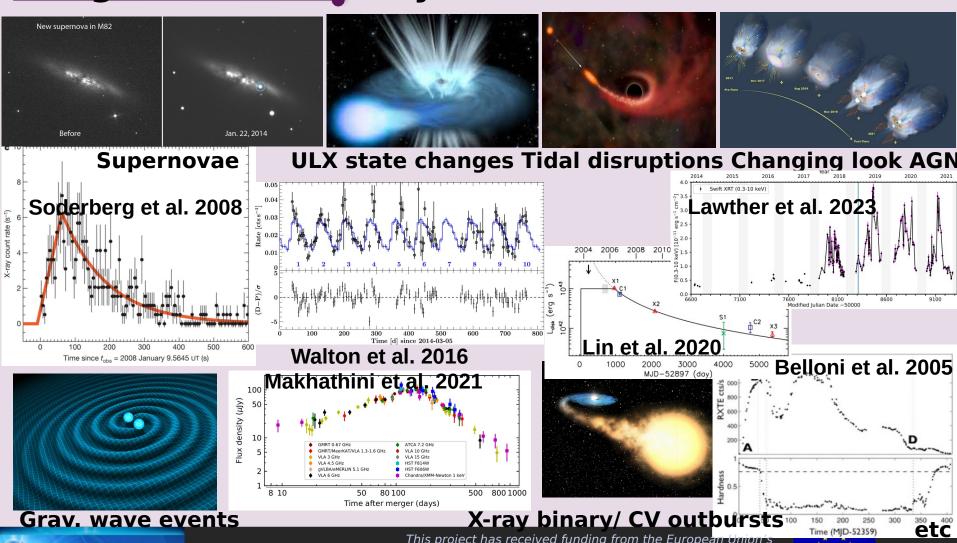






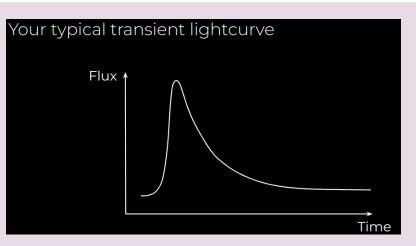
INTRODUCTION

Long-term variability

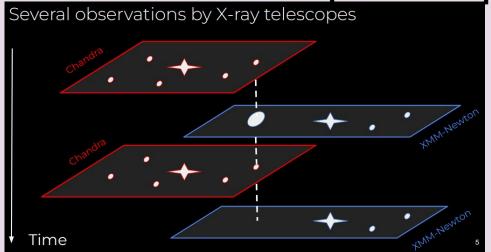


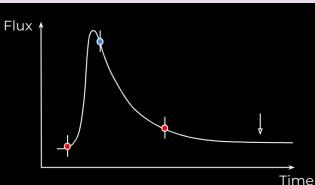


LONG TERM VARIABILITY



- Code developed to identify longterm variability (Quintin et al. 2024)
- Identifies factor >5 variation in flux
- Expected to send to screener to validate source then push to public
- Code placed in ACDS pipeline Reco.
 2023-05-11/11 but not activated as no protocol in place for send to IRAP



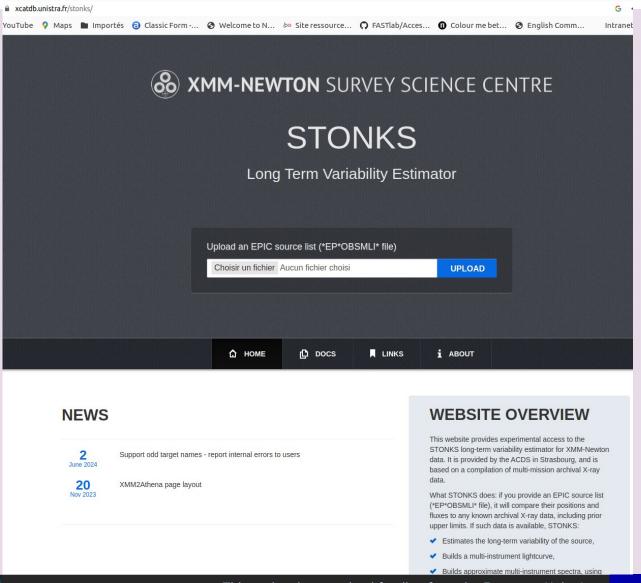








ONLINE INTERFACE





XMM Users Group

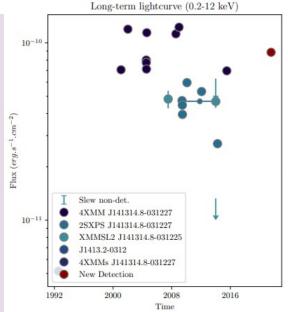
2023

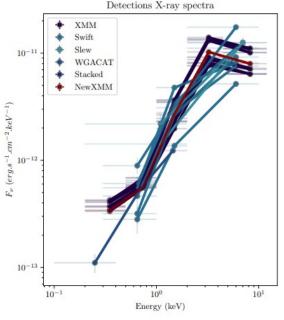
May 11th,

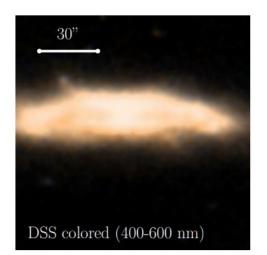


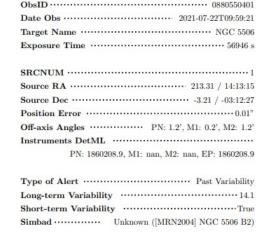


STONKS OUTPUT









!!! Warning: extreme spectrum might impact variability !!!

JSON file to push to web:

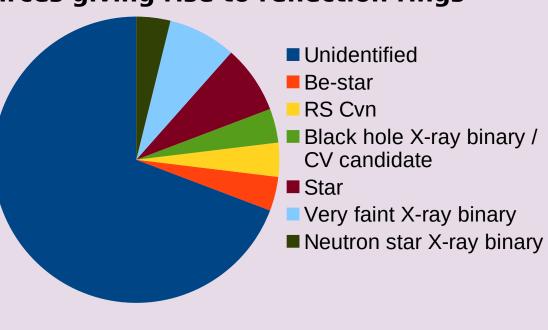
- > 0bsID
- > Date Obs
- > Exposure Time
- > SRCNUM
- > Source RA, Dec
- > Position Error
- > * DETML
- > * OFFAX
- > VarAmplitude
- > LastFlux, LastFluxErr
- > LastHR
- > ArchivalShortTermVar
- > Simbad id
- > SpectralWarning
- > AlertType:
 - 1. High Flux State
 - 2. Low Flux State
 - 3. Past Variability:
 - 4. First detection





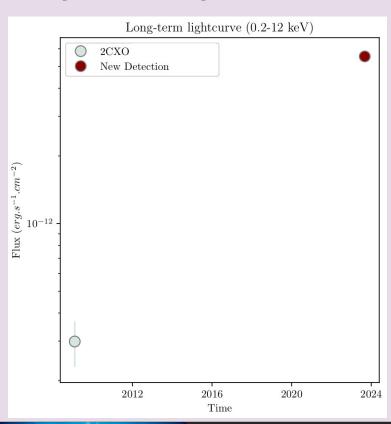
TESTING ALERTS WITH THE MULTI-YEAR HERITAGE PROGRAMME (MYHP)

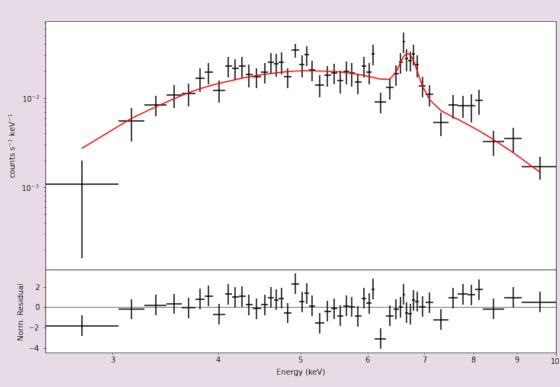
- Considered data from July 2023 onwards from Gabriele Ponti MYHP Galactic plane survey (only programme currently available)
- 87 observations, 4 with no detections
- 67 alerts raised
- 38 alerts were false detections mostly due Galactic plane containing many bright sources giving rise to reflection rings
- Of the 29 good alerts :
 - First ever detection: 14
 - High state: 7
 - Low state: 7
 - Past variability: 1
- 4 showed extreme spectra
- Only 3 sources in Simbad (stars)





- L ~3x10³⁴ erg s⁻¹ if at Galactic centre (given absorption and no optical counterpart)
- Possible very faint X-ray binary (Wijnands et al. 2006) small / symbiotic system

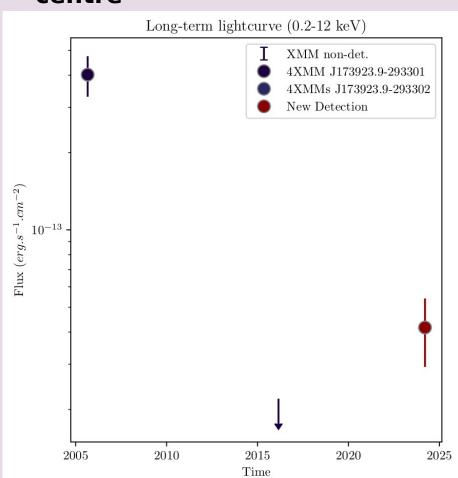


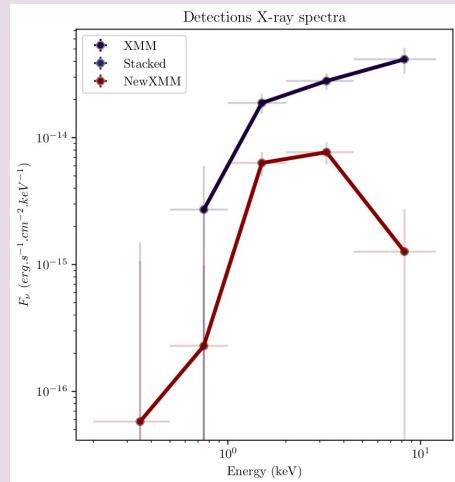






Candidate neutron star very faint X-ray binary close to Galactic centre

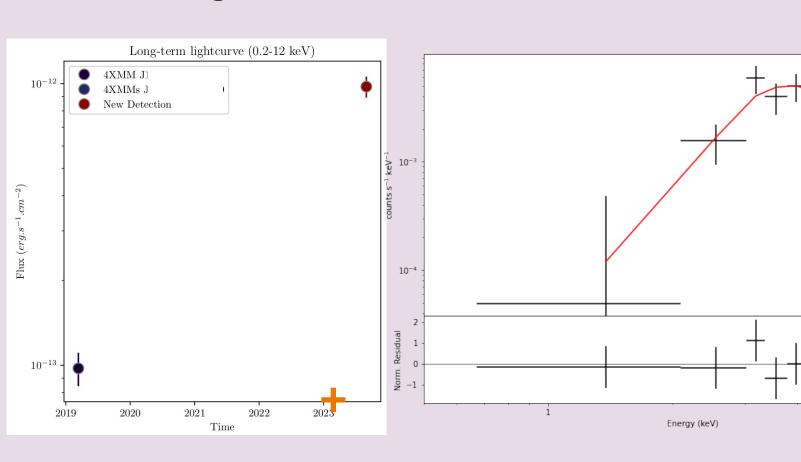








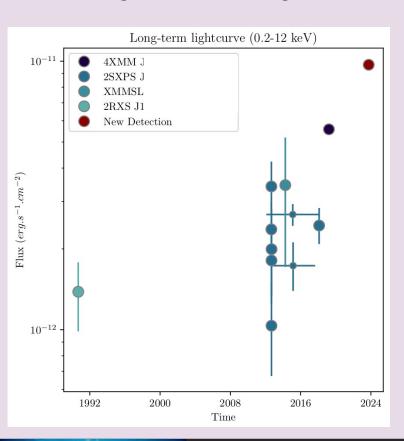
- New black hole X-ray binary/ cataclysmic variable
- L ~8x10³³ erg s⁻¹ if at Galactic centre

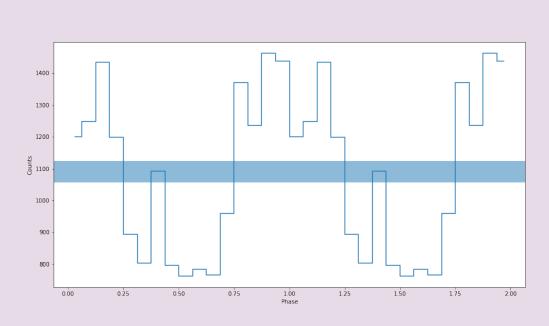






- New Be X-ray binary @1.8 kpc, 7980 s period, pulsed fraction: 0.29
- L ~4x10³³ erg s⁻¹ persistant system showing variability
- Hard power law spectrum ($\Gamma \sim 1.4$) + line @ 6.7 keV ($\sigma \sim 0.33$)









NEXT STEPS

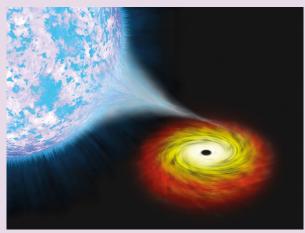
- New extra-galactic MYHP starting should allow longer term transients such as TDEs, supernovae, ULXs, changing look AGN, to be detected (and followed up), but also gravitational wave counterparts etc
- Screening with EPIC data essential for weeding out problems (screeners?)
- If not possible, protocol to be developed for sending to SOC & IRAP
- Web server developed at IRAP to go live to house alerts and send to interested parties
- Modify phase I (or II) to allow PI to uncheck a box when responding to AO if not happy about basic data made public about serendipitous transient







DISCOVERIES OF RARE SOURCES WITH THE SOFTWARE



- Origin of the very high luminosity observed from ultra luminous X-ray (ULX) sources unclear
- 7 show accelerating pulsations => neutron star (NS) compact object
- 8th candidate NS ULX found in galaxy
 with another NS ULX (Quintin et al.
 2021)
- Supports idea that many ULXs may host NS, implying emission is beamed and generated through fan beam geometry (Gnedin & Sunyaev, 1973)



- Quasi periodic eruptions (QPEs) discovered from massive black holes (Miniutti et al. 2019)
- Five systems known, two associated with tidal disruption events (TDEs)
- New good candidate found (Quintin et al. 2023), associated with TDE
- Suggests TDEs may be at the origin of QPEs and data gives constraints on the time from TDE to QPE
- Data helps understand the form of the eruption profile

