

Space physics made easy

An overview for the non developers

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How do you get in-situ plasma physics measurements?

Issues when the SciQLop project started (2014)

- Impossible to get all data from a single server (CDA, AMDA, CSA, CLWeb, PDS,...)
- Many file formats (TAB, CEF, CSV, CSV, CDF, netCDF,...)
- Relative compliance to standards (IST.. what?)
- Few Python packages, usually for one mission or one server
- Not ready for ML pipelines



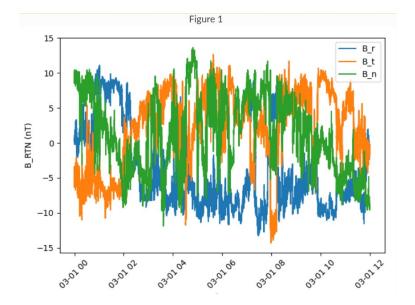
As easy as:

```
my_data = spz.get_data( what )
my_data = spz.get_data( what, start_time, stop_time )
```

Design principles:

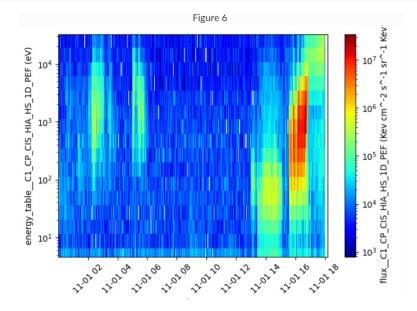
- Do not ever change the data
- Easy to Use Hard to Misuse (EUHM)
- Keep It Simple, Stupid (KISS)
- Separation Of Concerns (SOC)
- Avoid global states (except for cache)
- Performances matter

```
solo_fgm: SpeasyVariable = spz.get_data(
    spz.inventories.tree.cda.Solar_Orbiter.SOLO.MAG.SOLO_L2_MAG_RTN_NORMAL.B_RTN,
    "2022-03-01",
    "2022-03-01T12",
)
plt.figure()
solo_fgm.plot()
plt.tight_layout()
plt.show()
```





```
plt.figure()
spz.get_data(
    spz.inventories.tree.csa.Cluster.Cluster_1.CIS_HIA1.C1_CP_CIS_HIA_HS_1D_PEF.flux__C1_CP
    "2006-11-01",
    "2006-11-02",
).plot(cmap="jet")
plt.tight_layout()
plt.show()
```







You can also:

```
my_datas = spz.get_data( [what,...] )
```

```
my_datas = spz.get_data( [what,...], [(start, stop),...] )
```

```
products = [
    spz.inventories.tree.amda.Parameters.Wind.SWE.wnd_swe_kp.wnd_swe_vth,
    spz.inventories.tree.amda.Parameters.Wind.SWE.wnd_swe_kp.wnd_swe_pdyn,
    spz.inventories.tree.amda.Parameters.Wind.SWE.wnd_swe_kp.wnd_swe_n,
    spz.inventories.tree.cda.Wind.WIND.MFI.WI H2 MFI.BGSE,
    spz.inventories.tree.ssc.Trajectories.wind,
data_several_dates: List[List[SpeasyVariable]] = spz.get_data(
    products,
    spz.inventories.tree.amda.TimeTables.SharedTimeTables.SOLAR_WIND.Magnetic_Clouds
for i in range(5):
   #fig = plt.figure(figsize=(20, 6))
   fig = plt.figure()
   gs = fig.add_gridspec(5, hspace=0)
   axes = gs.subplots(sharex=True, sharey=False)
```

for j in range(5):

plt.tight_layout()

plt.show()

data_several_dates[j][i].plot(ax=axes[j])



Data discovery: an up to date hierarchical inventory

```
Jupyter console 6.6.3
Python 3.12.6 (main, Sep 9 2024, 00:00:00) [GCC 14.2.1 20240801 (Red Hat 14.2.1-1)]
Type 'copyright', 'credits' or 'license' for more information
IPython 8.25.0 -- An enhanced Interactive Python. Type '?' for help.
   spz.get_data(spz.inventories.tree.
```



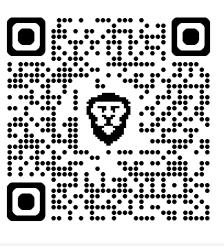
Speasy variables implements "Numpy" hooks (Numpy's Universal Functions)

```
def vect_to_mfa_delta(start, stop):
    B=spz.get_data(spz.inventories.tree.amda.Parameters.Cassini.MAG.orbit_saturn.cass_mag_krtphr.cass_b_krtphr,
                   start-timedelta(seconds=500),
                   stop+timedelta(seconds=500))
    B0=np.median(B, axis=0)
    delta B=B-B0
    # REF = [er, etheta, ephi]
    matrix_REF_to_MFA=np.empty((3,3))
    matrix_REF_to_MFA[0][:]=B0/np.linalq.norm(B0)
    matrix_REF_to_MFA[1]=np.cross(matrix_REF_to_MFA[0],[1,0,0])
    matrix_REF_to_MFA[2]=np.cross(matrix_REF_to_MFA[0],matrix_REF_to_MFA[1])
    B_mfa = np.einsum('kj,ij->ik', matrix_REF_to_MFA, delta_B)
    B_mfa = sosfiltfilt( sos=sos,var=B_mfa)
    return B mfa
```



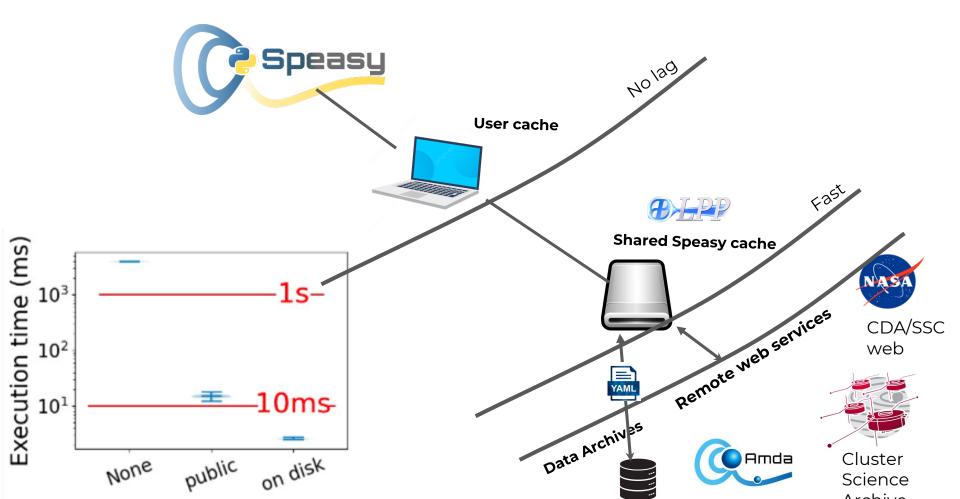
Scipy compatibility

- Resampling
- Interpolating a list of variables onto a reference one
- Filtering signal



```
def mirror mode threshold(start time: float, stop time: float) -> SpeasyVariable or None:
    mms1 products = spz.inventories.data tree.cda.MMS.MMS1
    products = [mms1 products.DIS.MMS1 FPI FAST L2 DIS MOMS.mms1 dis temppara fast,
                mms1 products.DIS.MMS1 FPI FAST L2 DIS MOMS.mms1 dis tempperp fast,
               mms1 products.FGM.MMS1 FGM SRVY L2.mms1 fgm b gse srvy l2,
                mms1 products.DIS.MMS1 FPI FAST L2 DIS MOMS.mms1 dis numberdensity fast]
    tpara, tperp, b, n = spz.get data(products, start time, stop time)
    anisotropy = tperp / tpara
    Pperp = tperp * n * 1e6
    b = interpolate(tperp, b)
    betaperp = Pperp * cst.mu 0 * cst.e * 2 / (b["Bt"] * 1e-9) ** 2
```

Super fast access to data



A custom CDF codec, CDF++/PyCDF++

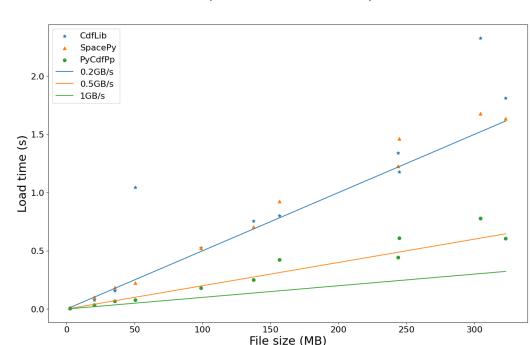


Designed to solve those issues:

- Thread safety
- Performances up to Python
- Fast and exact time conversion
- Use of modern C++
- Decouple logic from file format
- In memory files
- GPL license
- Can be used as testbed for CDF evolution or compression standards

POC of WebAssembly build;)

CDF implementations read speed



Access any local or remote archive

A "simple" YAML file to describe your archive:

```
mms1_scm_srvy_l1b_scsrvy:
     inventory_path: mammouth/MMS/MMS1/scm/scsrvy
    master cdf: >-
            /mammouth/data/mms1/scm/srvy/llb/scsrvy/2017/01/mms1_scm_srvy_llb_scsrvy_20170101_v1.1.0.cdf
     split_frequency: daily
     split_rule: regular
    url_pattern: >-
            /mammouth/data/mms1/scm/srvy/l1b/scsrvy/\{Y\}/\{M:02d\}/mms1\_scm\_srvy\_l1b\_scsrvy\_\{Y\}\{M:02d\}_v \ d+. \ d+
    use file list: true
mms1_edp_brst_l2_hmfe:
    fname\_regex: 'mms1\_edp\_brst\_l2\_hmfe\_(?P < start > d+)\_v(?P < version > \lceil d \setminus . \rceil +) \setminus .cdf'
     inventory_path: cda/MMS/MMS1/EDP/BURST
    master_cdf: >-
           https://cdaweb.gsfc.nasa.gov/pub/software/cdawlib/0MASTERS/mms1_edp_brst_l2_hmfe_00000000_v01.cdf
     split frequency: monthly
     split rule: random
    url pattern: >-
           use file list: true
```



- Easy to use (spz.get_data(product, time range))
- ~70k products "out of the box" from 4 different servers (more to come)
- Actively developed and maintained
 - HAPI client soon (I swear!)
 - Custom user codecs
- Already used for publications
- Plays well with Numpy and SciPy
- Will be used as AMDA gateway to CDAWeb